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Harmonic forcing in rotating flows: zonal flow, inertial waves, instabilities

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Geo-/Astrophysical context





Glatzmaier & Roberts, Nature **377** (1995)



Mechanical forcing



Gravitational effects





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1

Applications



Investigation of internal structure (Europa Jupiter System Mission : exploration of Jupiter's moons) Generation of zonal winds in atmosphere and core (alternative mechanism for patterns, dynamo,...)



Model : a rotating incompressible fluid enclosed in a container Study of the generated flow Evaluate the role of each process

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Image Credits: NASA/JPL

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Longitudinal and latitudinal libration

Rotation of a satellite at the mean angular velocity Ω_0

Longitudinal libration: small oscillation of the rotation rate of container $\Omega(t) = \Omega_0 + \delta\Omega \, \cos(\omega \, t)$

Latitudinal libration: small oscillation of the rotation axis of container





Harmonic Forcings

Animation Credits: NASA/JPL



Small amplitudes \longrightarrow Negligible forcings?

But can drive inertial waves & instabilities...

Fluid flows driven by harmonic forcing ?

Outline

- 1 Theory of rotating flows
- 2 Methods of investigations
- 3 Mean zonal flow
- 4 Resonance of inertial modes
- 5 Spontaneous generation of inertial waves



Outline

1 - Theory of rotating flows

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Navier-Stokes equation

Navier-Stokes equation in the rotating frame at $\Omega=\Omega_0\,{\bf k}$

$$\frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla)\mathbf{u} + 2\,\Omega_0 \,\times\, \mathbf{u} = -\frac{\nabla P}{\rho} + \nu\,\Delta\mathbf{u}$$

Timescale: Ω_0^{-1} Lengthscale: RVelocity scale: $U \longrightarrow associated with$ the forcing (ε, β)

Dimensionless form:



From planets to laboratory



From planets to laboratory

Amplitude of the forcing $R_0 \sim \epsilon, \ \beta$		
	Libration	Tides
Planets	$arepsilon \sim 10^{-4} - 10^{-3}$	$eta \sim 10^{-6} - 10^{-4}$
Laboratory	$arepsilon \sim 0.02 - 0.1$	$eta \sim 0.02 - 0.08$
Frequency of the forcing Libration Tides		
Planets	$\omega \sim 10^{-3}-1$	$\Omega_R \sim 0-2$
Laboratory	$\omega \sim 10^{-2}-1$	$\Omega_R \sim 0-1$
Ekman number $E = \frac{v}{\Omega_0 R^2}$		
Planets	$E \sim 10^{-15} - 10^{-12}$	
Laboratory	$E\sim 10^{-6}$ $-$	10^{-4}



Inertial waves

$$\frac{\partial \mathbf{u}}{\partial t} + R_o \left(\mathbf{u} \cdot \nabla \right) \mathbf{u} = -\nabla p + E \Delta \mathbf{u} - 2 \mathbf{k} \times \mathbf{u}$$

Coriolis force: restoring force

Plane inertial waves $(\mathbf{v} \cdot \nabla)\mathbf{v} + 2\mathbf{\Omega} \times \mathbf{v} = -\nabla P + E\nabla^2 \mathbf{v}$ $\mathbf{u} = \mathbf{u}_0 e^{i(k_x x + k_y y + k_z z - \omega t)}$ Dispersion inertial waves for $\omega = 2 \frac{k_z}{|\mathbf{k}|} = 2 \cos \gamma$ No inertial waves for $\omega > 2$



 $m{v} = m{v_0} \, \mathrm{e}^{\mathrm{r}} (\hat{m{u}}_t + \mathcal{W}_h) \, \hat{m{v}}_{\mathcal{P}}$ are usually damped by viscosity but can be forced...

 ω : frequency

m. Appin Sautital waven im Harmonic forcing in rotating flows



Inertial modes

Bounded volume (sphere, cylinder)

wave becomes phase-coherent, multiple reflection : **resonance**

inertial modes







Boisson et al., Phys. Fluids, 24 (2012)



Key features





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1 - Theory of rotating flows

2 - Methods of investigations

- 3 Mean zonal flow
- 4 Resonance of inertial modes
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From planets to the laboratory



Experimental setup



From sphere to spherical shell





Addition of an inner core



Numerical method



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Experimental observations

Longitudinal libration in a sphere $\Omega(t) = 1 + \varepsilon \cos(\omega t)$ $E = \frac{v}{\Omega P^2}$



Mathematical formulation

Sauret et Le Dizès, JFM, (in press)



Geometry: spherical shell $a = \frac{R_i}{R_e}$

 R_e external radius R_i internal radius

Fluid: rotating, incompressible, homogeneous, viscous

- kinematic viscosity
- mean rotation rate

Longitudinal libration

 $E = \frac{\nu}{\Omega_0 R_e^2}$

- ω_l frequency of libration
- $\delta \Omega_e$ amplitude of libration of the outer sphere
- $\delta \Omega_i$ amplitude of libration of the inner sphere

$$\omega = \frac{\omega_l}{\Omega_0}$$

$$\varepsilon = \frac{\delta \Omega_e}{\Omega_0} \qquad \qquad \alpha = \frac{\delta \Omega_i}{\delta \Omega_e}$$

Framework of the asymptotic analysis



Harmonic flow structure



Sauret et Le Dizès, JFM, (in press)
Generated by the non-linear interactions
$$u_2 = \tilde{u}_2 e^{2i\omega t} + \bar{u}_2$$

In the bulk:
 $U_2 = 0$
 $U_1 + (U_1 \cdot \nabla) U_1 + (2k \times U_2 = -\nabla p_2) + E\Delta U_2$
 $= O(E) = O(1) = O(E)$
 $U_2 = 0$
 $U_2 = 0$
Taylor-Proudman theorem

Non linear interactions in the Ekman layers

Stationary flux in the bulk

Equilibrium of the flux: induced the mean zonal flow

Sauret et Le Dizès, JFM, (in press)



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4

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Solution in Bulk I
$$\mathcal{X}_2(\rho) = \mathcal{F}(\rho; \omega)$$

 $\Psi_2(\rho) = 0$

with
$$s_{\pm} = \sqrt{\frac{\omega}{2} \pm \cos \theta}$$

Stewartson layer structure

External Outer Layer ${\bf E}^{\,1/4}$

smoothing of azimuthal velocity

$$u_{\Phi_2} \sim rac{\mathcal{F}(a;\omega)}{a} - a \,\delta \,\Omega_2 \,\mathrm{e}^{- ilde{
ho}} \ ilde{
ho} = (
ho - a) / \left[(1 - a^2)^{3/8} \, E^{1/4}
ight]$$

Internal Outer Layer ${\bf E}^{\,2/7}$

smoothing of axial vorticity $\omega_{z_2} = E^{-1/4} \frac{\delta \Omega_2 a}{(1-a^2)^{3/8}}$ $\bar{\rho} = (a-\rho)[a/(2 E^2 (1-a^2)^2)]^{1/7}$ Stewartson, JFM, **26** (1966) Sauret et Le Dizès, JFM, (in press)



Inner Layer $E^{1/3}$

smoothing of axial velocity

$$u_{z_2} \sim E^{5/42} \, \frac{a_0 \, \delta \Omega_2}{a^{81/28} \, (1-a^2)^{25/84}} \, \, \mathbf{J}\left(\frac{z}{\sqrt{1-a^2}}, \frac{\rho-a}{E^{1/3} \, (1-a^2)^{1/6}}\right)$$

Sauret et Le Dizès, JFM, (in press)

Generated by the non-linear interactions $u_{\Phi_2} = \frac{\chi_2}{\rho} \qquad u_{z_2} = \frac{\sqrt{E}}{\rho} \frac{\partial \Psi_2}{\partial \rho} \qquad \omega_{z_2} = \frac{1}{\rho} \frac{\partial \chi_2}{\partial \rho}$

$$\mathbf{u}_2 = \mathbf{\tilde{u}}_2 \, \mathrm{e}^{2\,\mathrm{i}\,\omega\,t} + \mathbf{\bar{u}}_2$$

$$\begin{aligned} \mathcal{F}(\sin\theta;\omega) &= \frac{\sin^2\theta}{32} \left[8 + \frac{(l_+ + \sqrt{\cos\theta}) (4 l_+^2 \cos\theta + \sin^2\theta)}{l_+^3 (2 l_+^2 + 2 l_+ \sqrt{\cos\theta} + \cos\theta)} + \frac{(l_- + \sqrt{\cos\theta}) (4 l_-^2 \cos\theta + \sin^2\theta)}{l_-^3 (2 l_-^2 + 2 l_- \sqrt{\cos\theta} + \cos\theta)} \right. \\ &- \frac{4 (l_-^3 + l_-^2 \sqrt{\cos\theta} + (l_- + l_+) \cos\theta}{l_+ l_- (l_- + \sqrt{\cos\theta})} + \frac{l_-^4 \sin^2\theta (l_- + \sqrt{\cos\theta} + (l_+^3 + l_-^3)/l_-^4 \cos\theta)}{\cos\theta l_+^3 l_-^3 (l_- + \sqrt{\cos\theta})} \\ &- \frac{4 (l_+^3 + l_+^2 \sqrt{\cos\theta} + (l_+ + l_-) \cos\theta}{l_- l_+ (l_+ + \sqrt{\cos\theta})} + \frac{l_+^4 \sin^2\theta (l_+ + \sqrt{\cos\theta} + (l_-^3 + l_+^3)/l_+^4 \cos\theta)}{\cos\theta l_-^3 l_+^3 (l_+ + \sqrt{\cos\theta})} \\ &+ \frac{1}{\cos\theta l_-^3 l_+^3 (l_- + l_+) (l_- + l_+ + \sqrt{\cos\theta})^2} \\ &\left(4 l_-^2 l_+^2 (l_- - l_+)^2 \cos^{5/2} \theta \left[1 + 2 (l_- + l_+) \sqrt{\cos\theta} \right] \\ &+ \cos\theta \left[l_- l_+ (l_- + l_+) (2 (2 l_-^2 - 3 l_- l_+ + 2 l_+^2) \sin^2\theta + 4 l_-^5 l_+ - 8 l_-^3 l_+^3 + 4 l_- l_+^5) \\ &+ 2 (l_-^5 + l_+^5) \sin^2\theta \right] + (l_- + l_+)^2 (l_-^4 + l_+^4 + l_- l_+ (l_- - l_+)^2) \sin^2\theta \right] \right]. \end{aligned}$$

$$U_{\pm} = \sqrt{\cos\theta \pm \frac{\omega}{2}}$$

Valid for $\omega < 2 \cos \theta$



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Sauret et Le Dizès, JFM, (in press)

22

Limit $\omega \rightarrow 0$



Deviation to solid-body rotation

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Comparison and scalings $\omega \rightarrow 0$



Comparison and scalings $\omega > 2$

Sauret et Le Dizès, JFM, (in press)





Spherical shell - bulk



Spherical shell - Stewartson layers



Analogy between azimuthal and time forcing

Sauret A. (submitted to Physica D) Libration (m=0) forces a Tides (m=2) forces a flow of azimuthal wavenumber flow of azimuthal wavenumber m=0m=2Solid body Harmonic flow rotation ϵ^2 $\varepsilon \mathbf{u}_1$ \mathbf{u}_0 \mathbf{u}_2 **u** = Double-harmonic flow Mean flow correction $e^{2i\omega t}$ $e^{-2i\omega t}$ e^{i ω t} $e^{2i\theta}$ $(\mathbf{u} \cdot \nabla) \mathbf{u}$ $(\mathbf{u}\cdot \nabla$ \mathbf{u} -4iθ e^{−iωt} $e^{-2i\theta}$ e^0 Mean zonal flow (axisymmetric & steady) Alban Sauret Harmonic forcing in rotating flows 01/02/2013 28

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Kalliroscope vizualisation



Solid-body rotation

$$E=9.46 imes10^{-6}$$
, $eta=0$, $\Omega_R=0$



Kalliroscope vizualisation

$$R_{ext} = 1 + rac{eta}{2} \cos(2\,\Phi)$$
 $E = rac{
u}{\Omega_0 R^2} \qquad \Omega_R = rac{\Omega_{orb}}{\Omega_0}$



 $E=1.4 imes10^{-5}$, eta=0.05, $\Omega_R=0.18$





Presence of a differential rotation

Influence of the frequency



Modes of the sphere



PIV Measurements



 $\Omega_R = 0.178$

$$\Omega_R = 0.384$$



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Influence of tides amplitude



Influence of the Ekman number



Shear instability



 $\Omega_R = 0.384, \ \beta = 0.045 \ \& \ E = 9.46 imes 10^{-6}$



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Shear instability





 10^{-1} $\Omega_{R} = 0.384$. * Unstable E^{0.71} 0 Stable 10⁻² 10⁻⁵ Ε

Critical deformation: $\beta_c \propto E^{0.71}$

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 $\Omega_R = 0.384$





Critical deformation: $\beta_c \propto E^{0.84}$

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Inviscid criterion:
$$\sigma_{NV} \sim \omega_z \sim \frac{\partial U_{\Phi}}{\partial \rho}$$
 $\omega_z = \frac{U_{\Phi}}{\rho} + \frac{\partial U_{\Phi}}{\partial \rho}$

Viscous damping term: $-k^2 E$

Thickness of the shear layer: $ho -
ho_c \sim E^{1/5}\, ilde
ho$

Noir et al., JFM **437** (2001) Kida, JFM 680 (2011)

Experimental results: $U_{\Phi} \sim \rho + \beta^2 E^{-a} \Omega_2(\rho)$



Inviscid criterion:
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 $\omega_z = \frac{U_{\Phi}}{\rho} + \frac{\partial U_{\Phi}}{\partial \rho}$

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Presence of an inner core



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Centrifugal instability



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Centrifugal instability

Longitudinal libration: $\Omega(t) = 1 + \varepsilon \cos(\omega t)$



Analytical determination of the threshold

Experiments (in a cylinder): at the threshold the instability is **axisymmetric**

Oscillatory flow (Stokes layer):

$$u_{1,\theta} = \varepsilon \exp\left(-\frac{1-r}{\sqrt{E}}\sqrt{\frac{\omega}{2}}\right) \cos\left(\omega t - \frac{1-r}{\sqrt{E}}\sqrt{\frac{\omega}{2}}\right)$$



Unstable at some instants

Centrifugal instability

Frozen state

Non viscous analysis

 \rightarrow Rayleigh criterion

$$\longrightarrow \Phi = \frac{\mathsf{d}}{\mathsf{d}r} \left(r^2 \, u_{\theta}^{(tot)^2} \right) < 0$$

Viscous analysis

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Sauret et al., Phys. Fluids, 24 (2012)

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Threshold

Experiments (in a cylinder): at the threshold the instability is axisymmetric

Oscillatory flow (Stokes layer):

$$u_{1,\theta} = \varepsilon \exp\left(-\frac{1-r}{\sqrt{E}}\sqrt{\frac{\omega}{2}}\right) \cos\left(\omega t - \frac{1-r}{\sqrt{E}}\sqrt{\frac{\omega}{2}}\right)$$

Centrifugal instability \longrightarrow Rayleigh criterion



Systematic study in a cylinder



Spontaneous generation of inertial waves



Preliminary mechanism

Localized source: $P_{init}(z, t) = f(z) g(t)$

Propagation:
$$p(x, z, t) = \frac{1}{4\pi^2} \int d\omega \int dk_z \hat{f}(k_z) \hat{g}(\omega) \exp\left[i(k_x x + k_z z + \omega t)\right]$$

Dispersion relation: $\omega = 2 \frac{k_z}{\sqrt{k_x^2 + k_z^2}} \longrightarrow k_x^2 = k_z^2 \left(\frac{4}{\omega^2} - 1\right)$

Evanescent waves: $|\omega|>2$ Propagative waves: $\omega\in[-2;2]$

$$p(x, z, t) \simeq \frac{1}{2 \pi} \int_{-2}^{2} d\omega f(z - x \sqrt{\frac{4}{\omega^{2}}} - 1) \hat{g}(\omega) e^{-i \omega t}$$

function localized
around $\omega = 1/\sqrt{1 + z^{2}/x^{2}}$
Spectrum of the
initial excitation

Selection of the frequency during the propagation Generic mechanism

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Focalisation & selection

 $\omega=2.1$, arepsilon=0.8, $E=4 imes10^{-5}$ u_r at r = 0.20.2 0.8 0.2 0.6 0.15 0.4 0.15 0.2 P(w) 0.1 Ν 0 0.1 -0.20.05 -0.40.05 -0.6 -0.8 2 3 4 4 3 2 0 1 ω ω **Propagation:** $\omega_{max}=rac{1}{\sqrt{1+rac{z^2}{(1-0.2)^2}}}$

Maximum radial flux of axial momentum

Dohan & Sutherland, Dyn. Atmos. Oceans, 40 (2005)

Least amount of viscous damping

Taylor & Sarkar, J. Fluid Mech., **590** (2007)



Conclusions

_____ Min: -5.00e-3

(3)



in a shell

Experimental and numerical validation

Analogy between azimuthal and time forcing Inertial waves



Experimental measurements of the **flow at resonance**

Characterization of the subsequent **shear instability**



Analytical study of the threshold and validation Numerical evidence of the spontaneous generation of inertial waves

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Perspectives



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Thank you for your attention